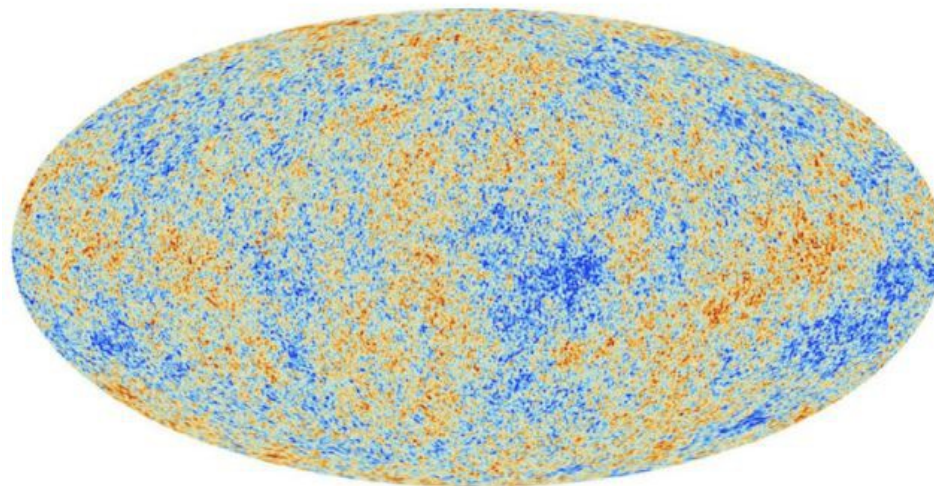
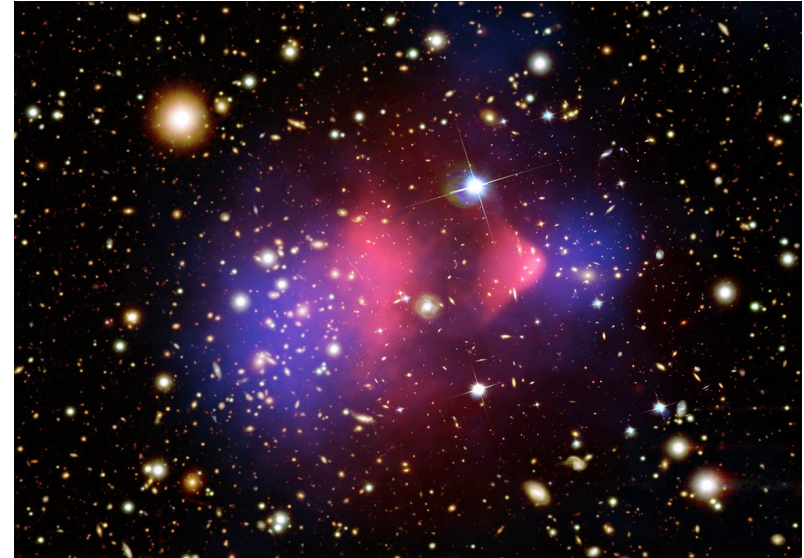
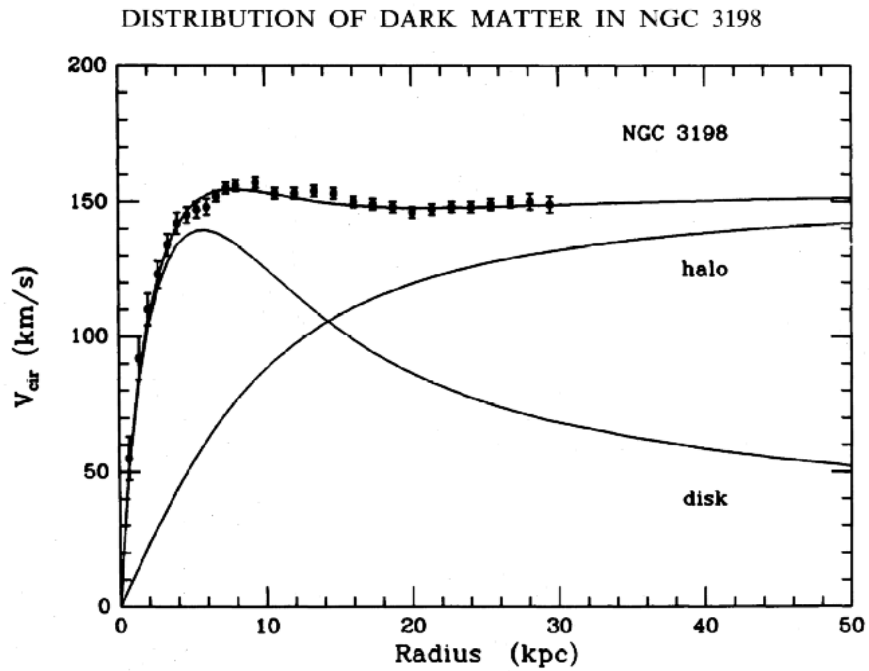


Complementarity of Dark Matter Searches at
Resonance (arXiv:1406.3288)

John Heal

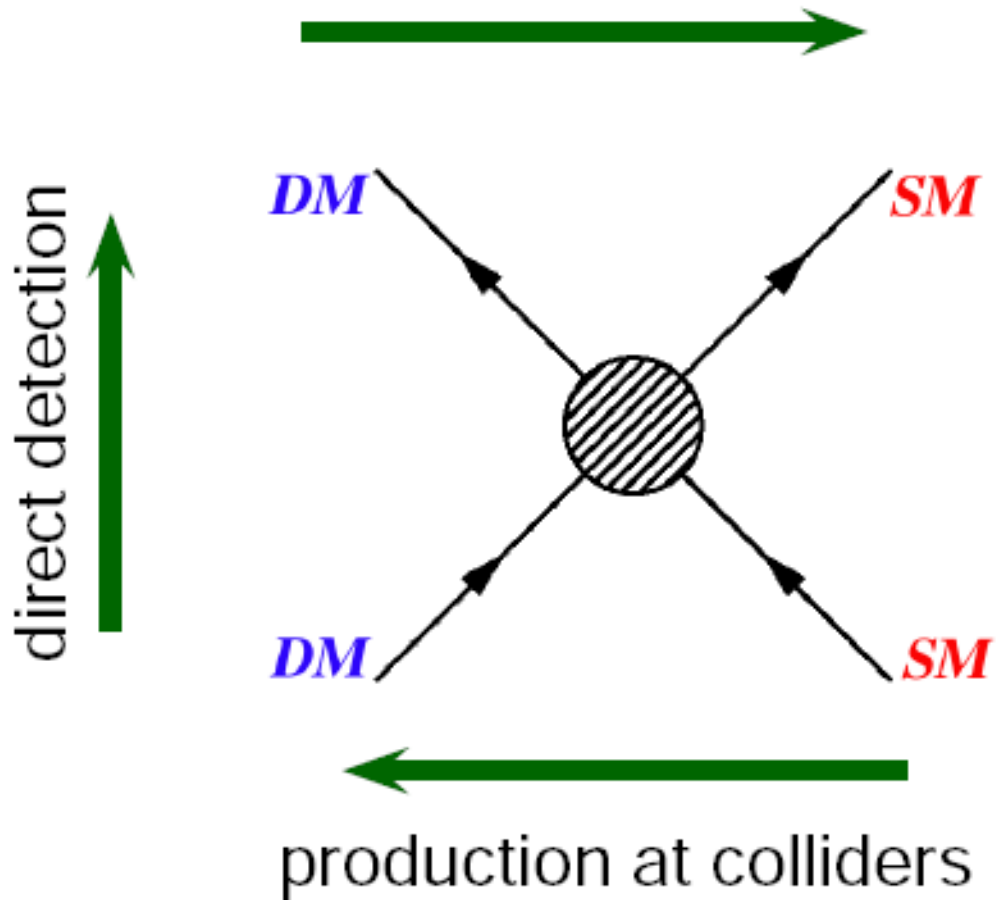
Supervisor: Malcolm Fairbairn

Dark Matter Evidence

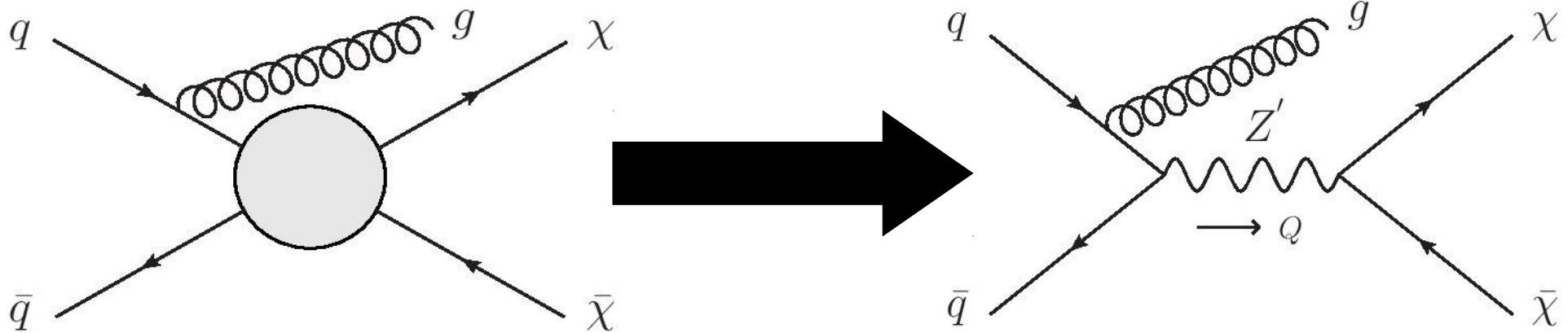


Effective Interaction

thermal freeze-out (early Univ.)
indirect detection (now)



Beyond Effective Interactions



Effective Interactions:

2 parameters.

Easier to constrain.

Very good for low energy physics.

Simplified Models:

More than 2 parameters.

More difficult to constrain.

Relationships between parameters can avoid direct detection limits.

Good for higher energy physics.

Open up new search channels.

Simplified Model

$$-\frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} + \frac{1}{2}m_{A'}^2 A'_\mu A'^\mu + \bar{\chi}(\gamma^\mu \partial_\mu - m_\chi)\chi \\ + A'_\mu \bar{\chi} \gamma^\mu (g_{\chi V} - g_{\chi A} \gamma^5)\chi + A'_\mu \bar{q} \gamma^\mu (g_{qV} - g_{qA} \gamma^5)q$$

Fermionic DM

Spin-1 Mediator

Leptophobic

Flavour blind coupling to quarks

Parameter ranges:

$$m_{A'} = 1.5 \text{ TeV}$$

$$m_\chi < 2.5 \text{ TeV}$$

$$0 < g_{\chi V}, g_{qV} \leq 3$$

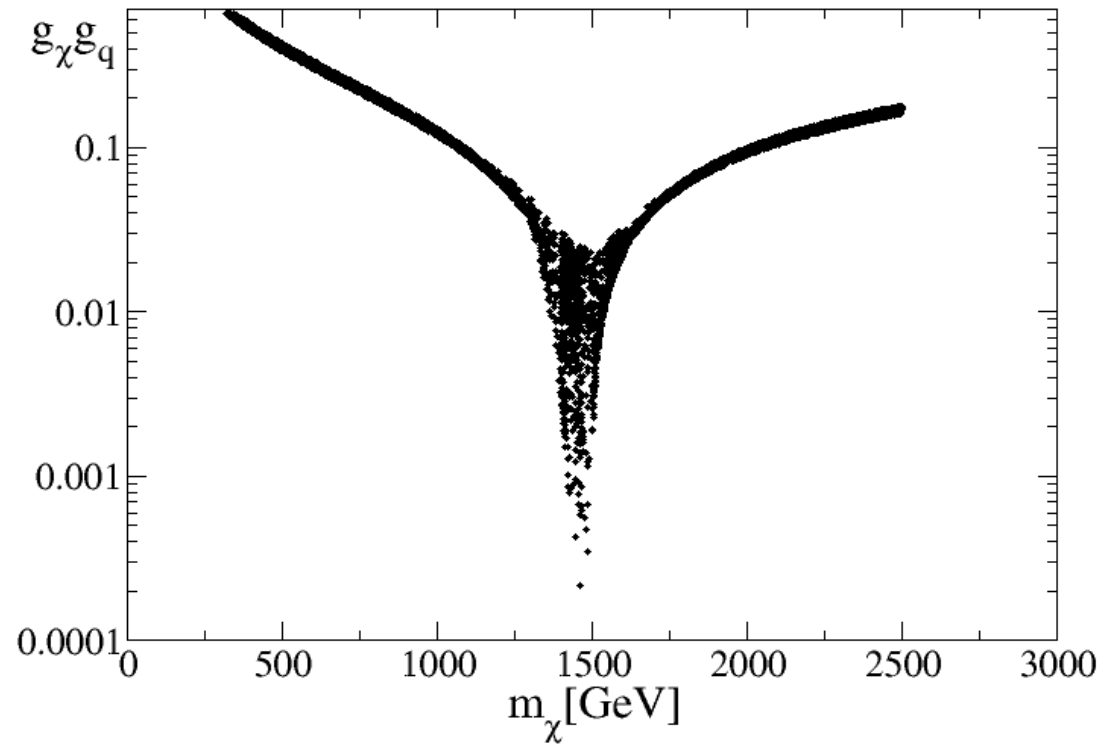
$$g_{\chi A}, g_{qA} = 0$$

$$\Gamma_{A'} \leq 0.15 m_{A'}$$

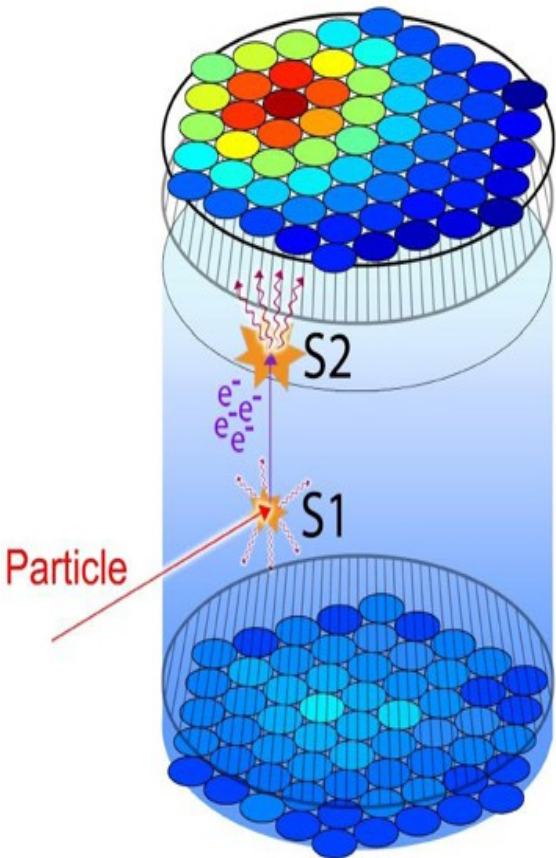
Relic Density

Assume standard freeze out
Calculate using micrOMEGAS
Getting from Planck:

$$\Omega_{DM}h^2 = 0.1198 \pm 0.0026$$



Direct Detection



- ionization electrons
- UV scintillation photons (~175 nm)

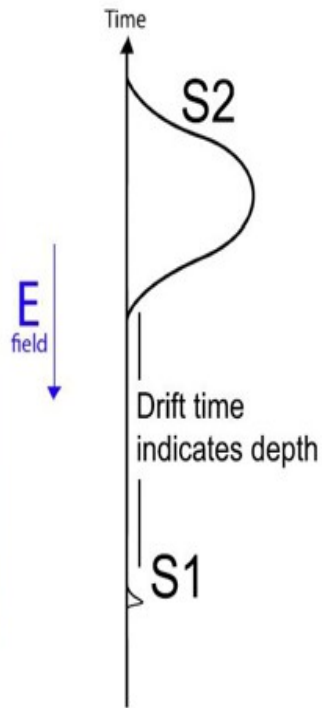
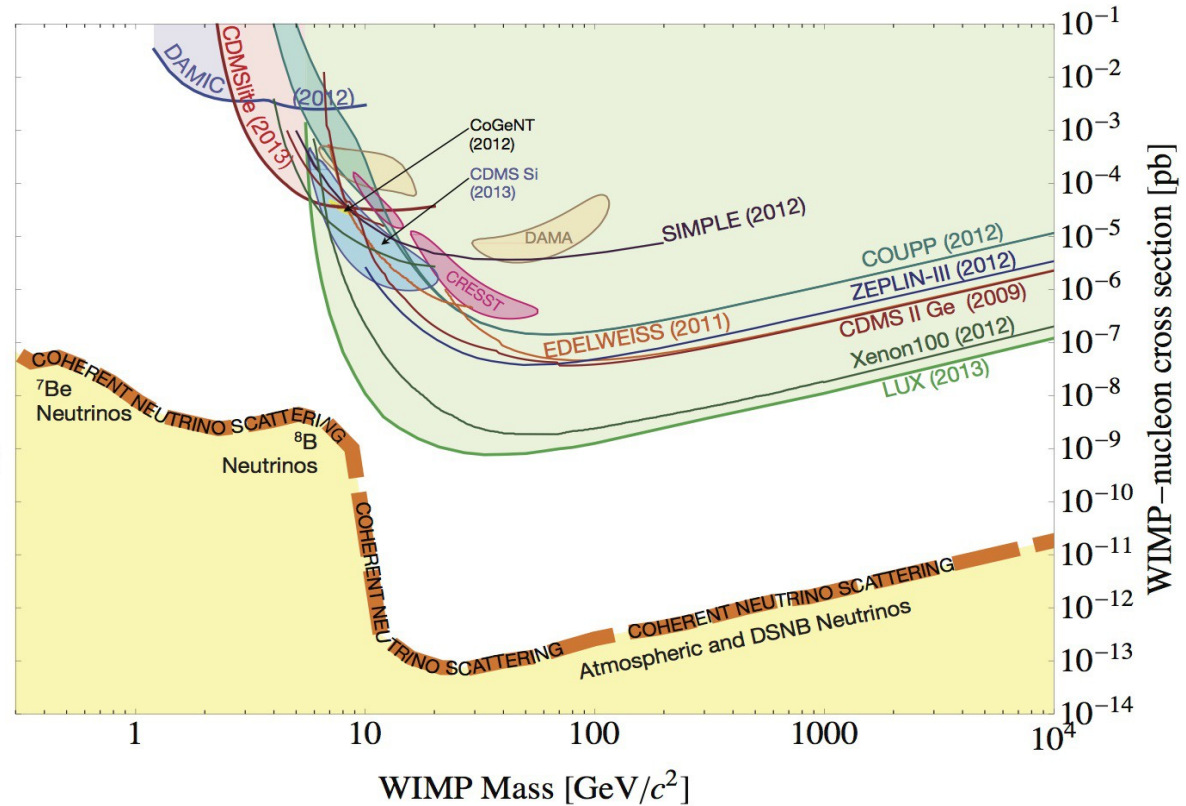
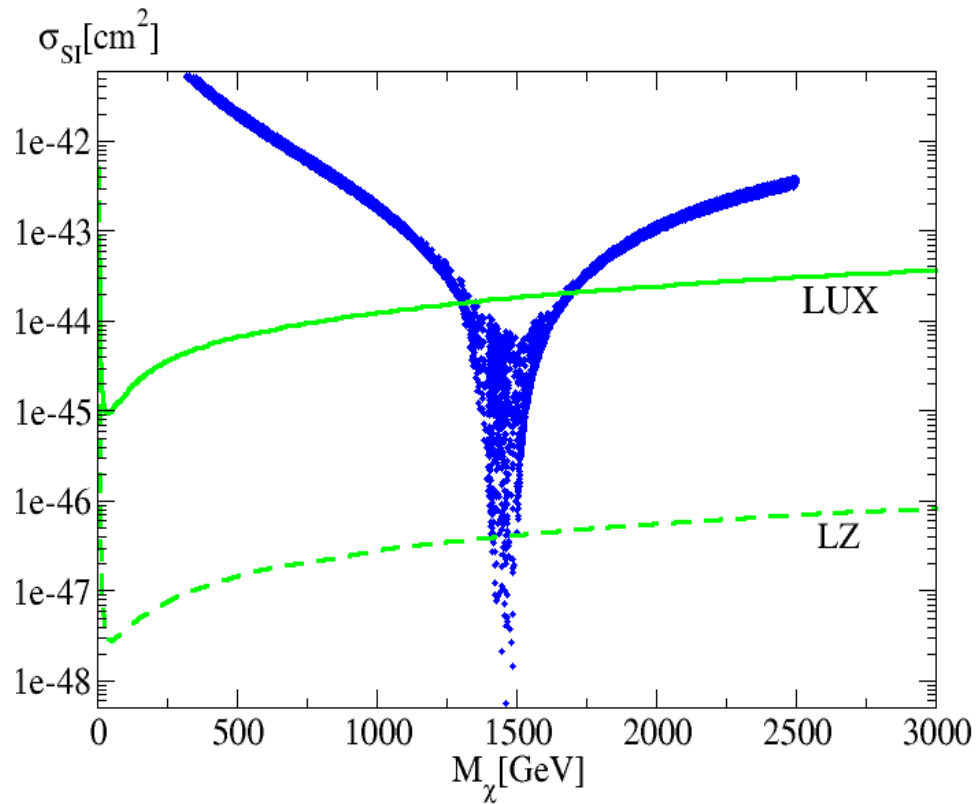


Image by CH Faham (Brown)



Direct Detection



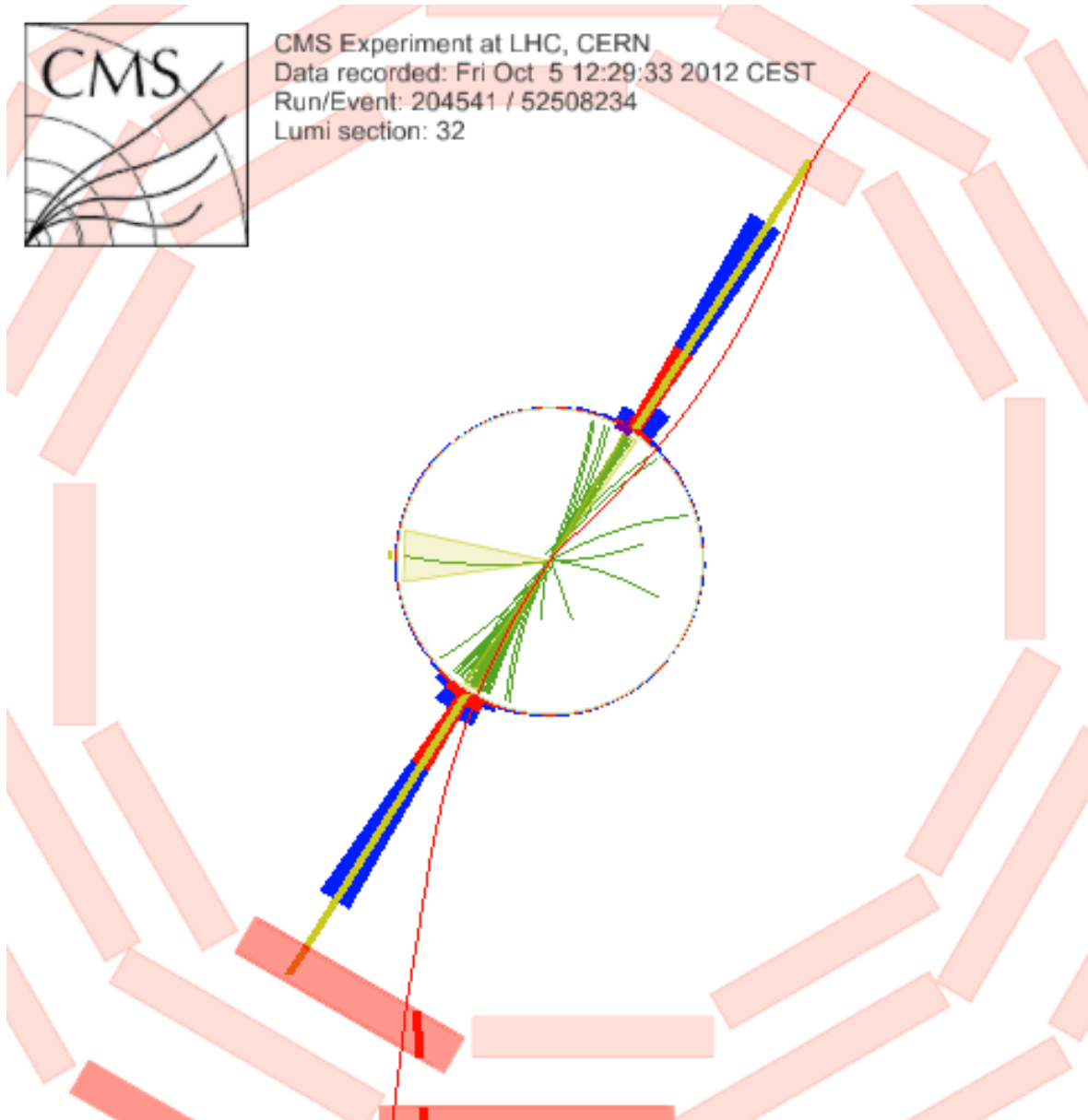
Characteristic shape due to the couplings falling off to produce the correct relic density when on the s-channel resonance

LUX limits the mass range to between roughly 1.25 and 1.75 TeV

Collider Searches - Dijets



CMS Experiment at LHC, CERN
Data recorded: Fri Oct 5 12:29:33 2012 CEST
Run/Event: 204541 / 52508234
Lumi section: 32



New heavy vector boson

Can possibly be produced on-shell and decay back into quarks

Experiments already looking at this signature

Result can be reinterpreted in the context of simplified models of dark matter

Not possible in the effective operator approach

Collider Searches – Dijet Validation

Steps:

Edit the model in Feynrules

Compute the matrix element in MadGraph

Generate and shower the events in pythia

Cluster the jets with FastJet

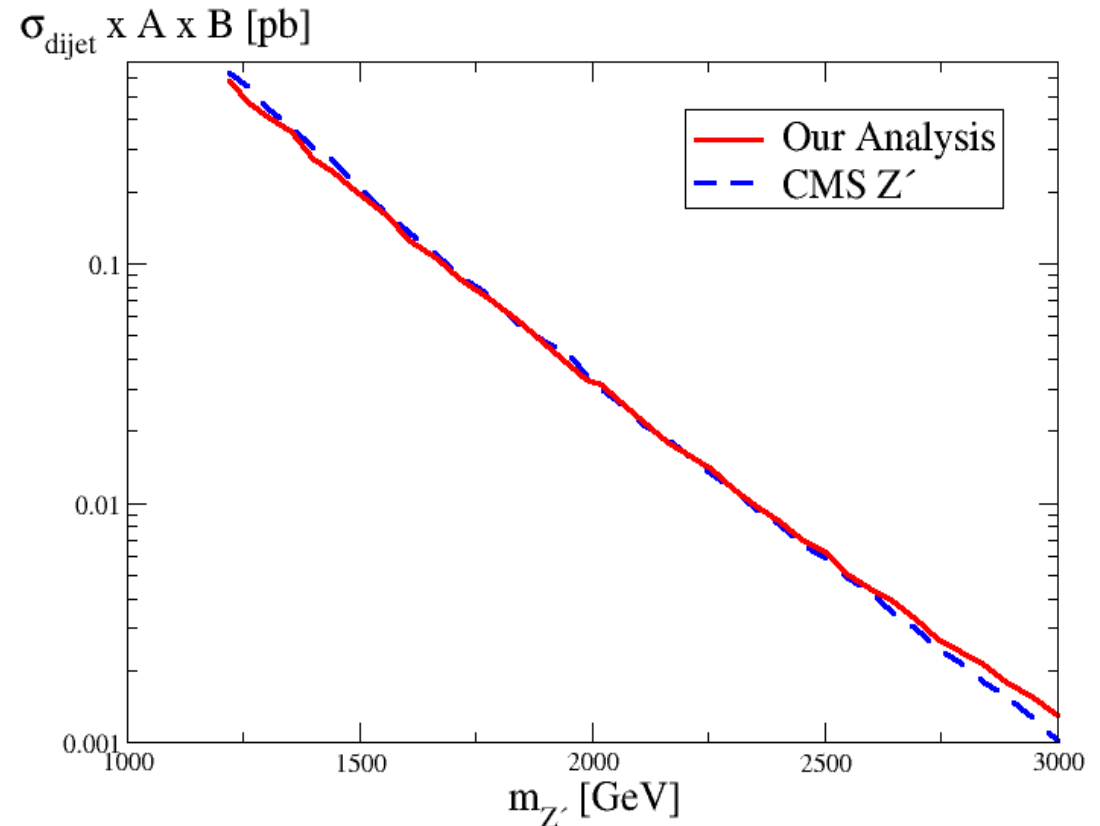
Apply CMS dijet cuts

$$p_T > 30 \text{ GeV}$$

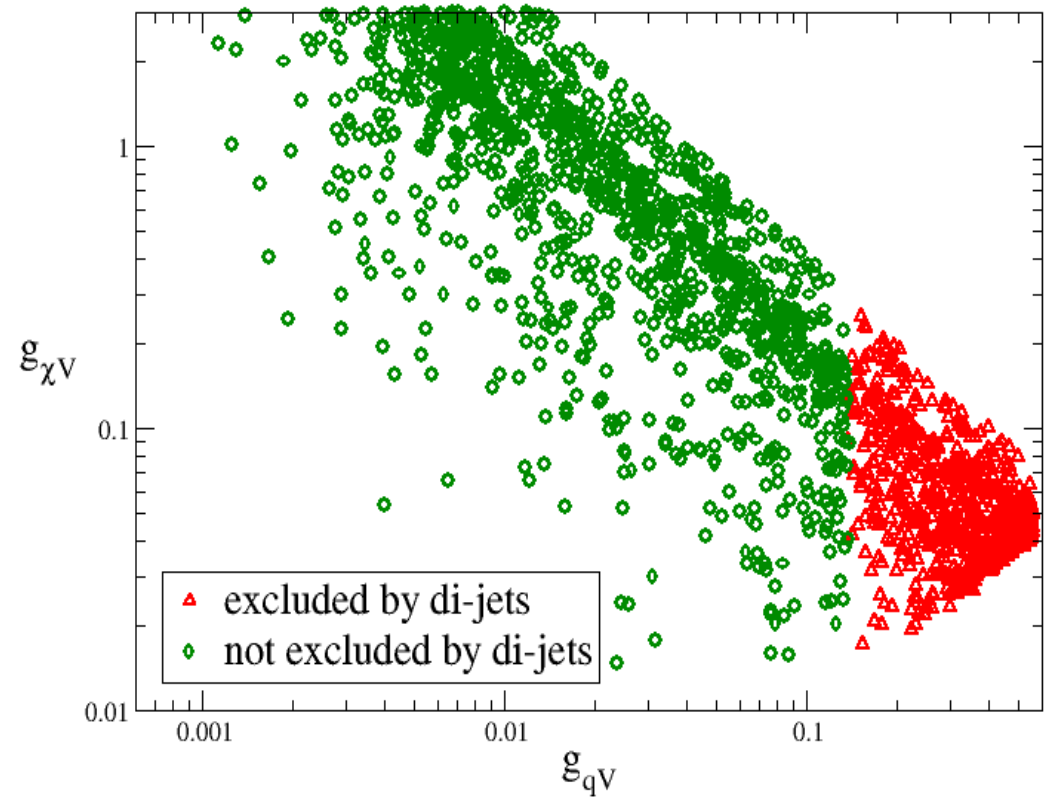
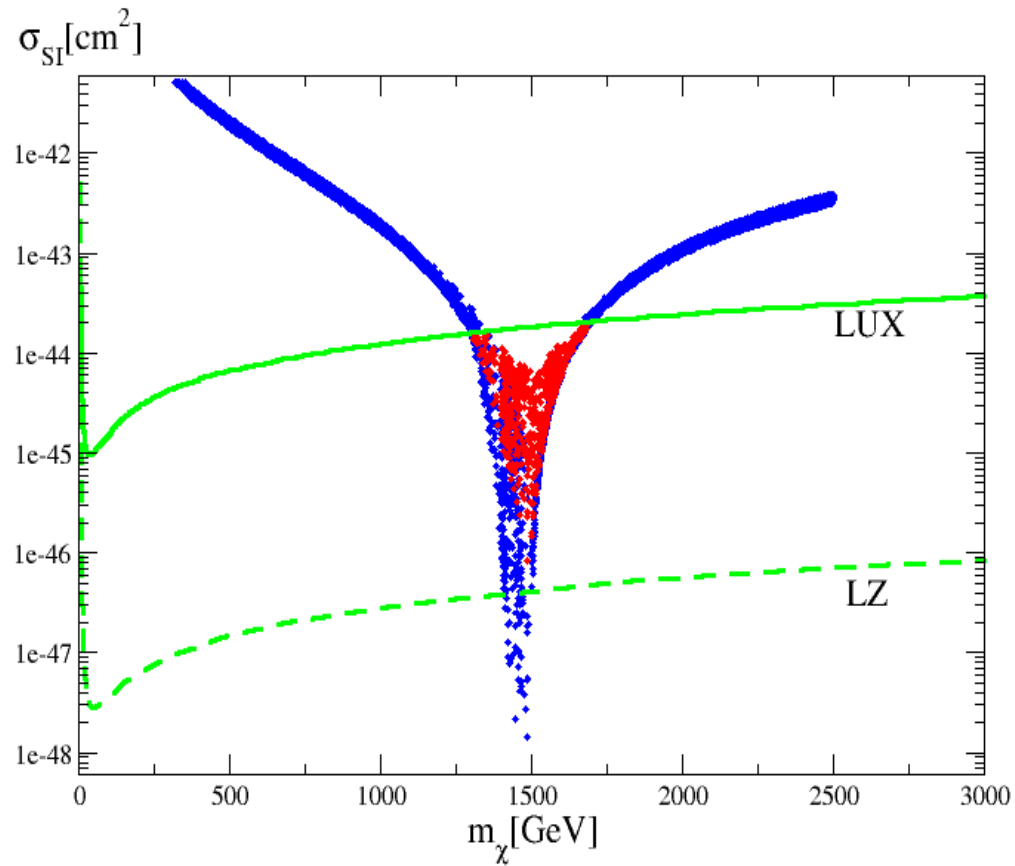
$$|\eta| < 2.5$$

$$m_{jj} > 890 \text{ GeV}$$

$$|\eta_{jj}| < 1.3$$



Collider Searches – Dijet Results

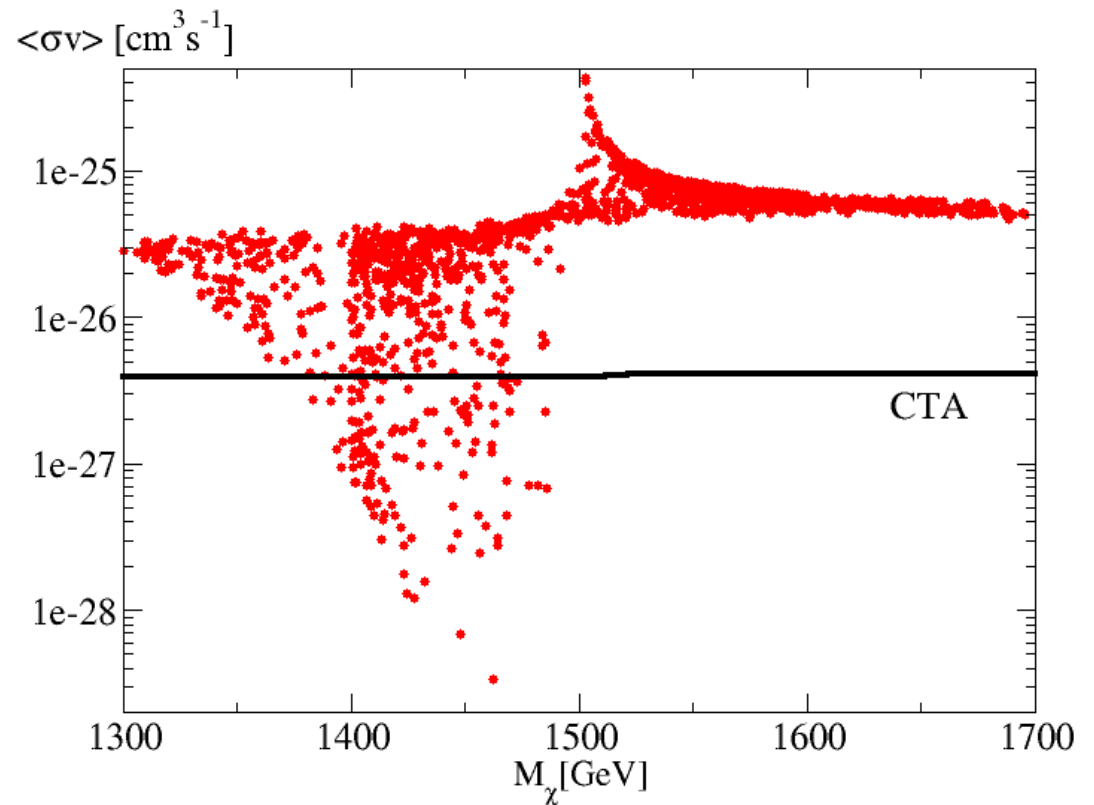


Indirect Detection

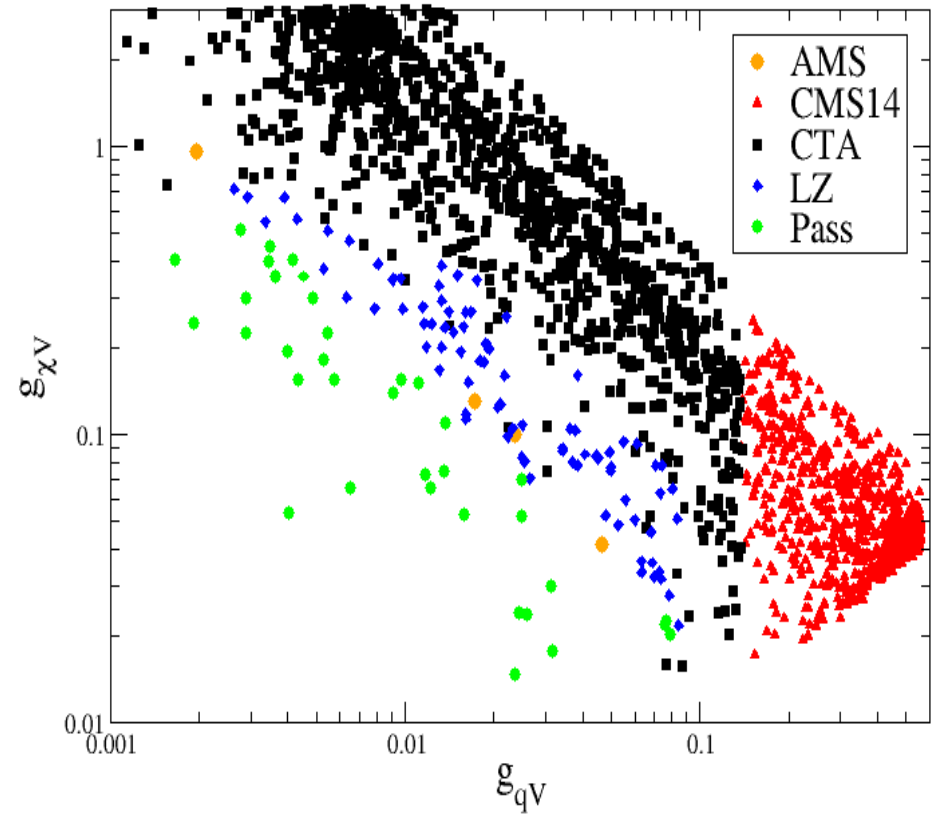
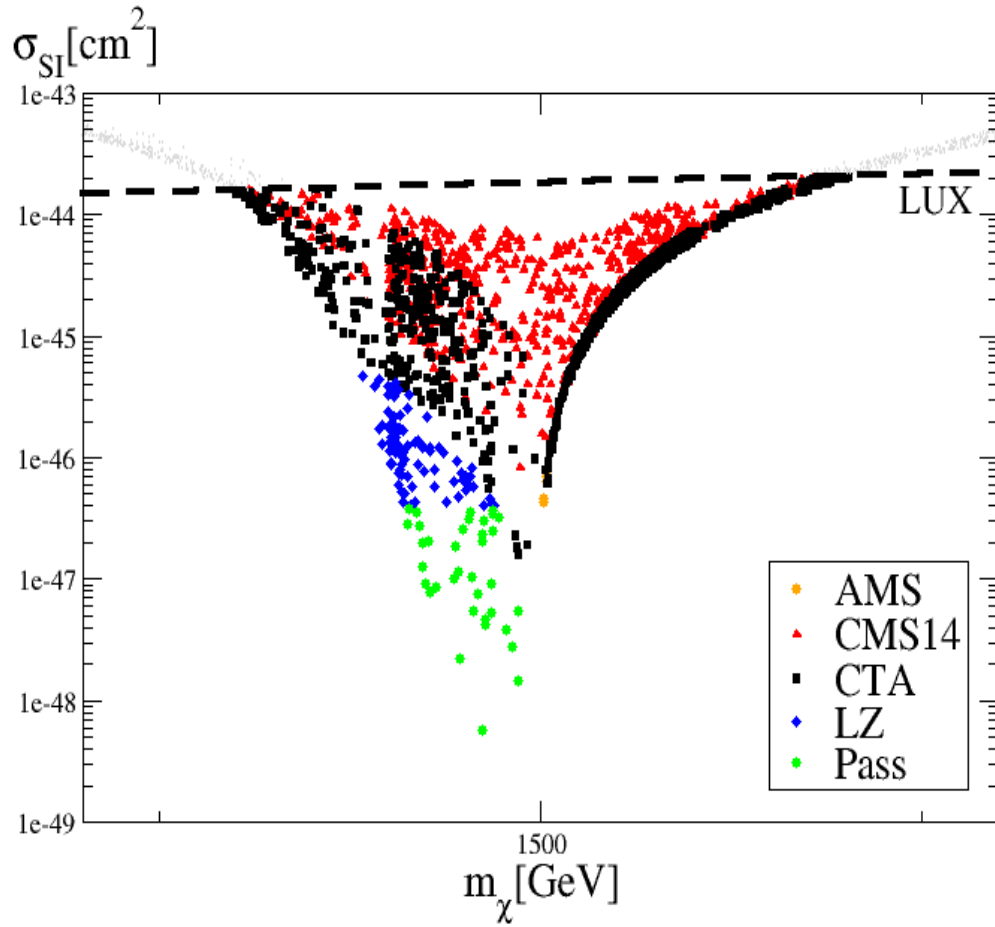
Self annihilation signals in regions of high dark matter density
Dwarf galaxies
Galactic centre

Same process as the relic density
with a different energy

Leads to a suppression of the cross section when DM is less than half the mass of the mediator



Combined Limits



Conclusions

Direct detection heavily constrains WIMP models of dark matter

These limits can be avoided in certain kinematic regions of parameter space

To study these simplified models should be used

These models open up new searches and possible constrains

Combining these searches can restrict a large amount of parameters